

# OPTICS AND TELESCOPES

## Chapter 6

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**INTRODUCTION** The telescope is the astronomer's window into the cosmos. The laws of optics govern how a telescope focuses incoming light to a focal plane. Smaller telescopes may be refracting or reflecting telescopes. A refracting telescope uses only lenses (refractive elements) to bring light to a focus while a reflecting telescope uses curved mirrors to bring light to a focus. Catadioptric telescopes use both lenses and mirrors. Refractors are more expensive and heavier than comparably sized reflecting telescopes. All professional observatory-sized telescopes are reflecting telescopes. Assuming that any telescope suitable for astronomical viewing is going to possess excellent optical components, the fundamental criteria for judging the quality of a telescope is its light-gathering power (LGP), which is determined by the aperture of the primary objective. Ultimately, regardless of aperture, the quality of the "seeing" will determine quality of the image produced by the telescope. Even the largest telescopes will produce blurred images if the atmosphere is not steady and clear. Adaptive optics, which are being used in most modern observatories, minimize or can even eliminate the effects of slow-moving atmospheric distortions in the telescopic image. Telescopes are not limited to viewing only visible light. High-altitude and orbiting observatories have been equipped with  $\gamma$ -ray, x-ray, UV, and IR telescopes. The high altitudes eliminate the absorption of the light due to the atmosphere. Ground-based telescopes may be optical or radio. Radio telescopes don't suffer as much from the atmosphere since radiowaves are only minimally affected by atmospheric conditions. However, the resolving power of a telescope is dependent on the observed wavelengths so radio telescopes must be very large (or combined as part of a long-baseline array) to achieve resolutions similar to optical telescopes.

- GOALS**
- ✓ It is important to know the different basic telescope designs and the different components which bring light to a focus.
  - ✓ Lab investigations showed how different optical components responded to light and how to measure the focal length of a lens. In lab, it was also demonstrated how to make a simple refracting telescope and calculate its magnifying power.
  - ✓ A basic knowledge of the different basic telescope mounts is useful when comparing different observatories.
  - ✓ It is important to know the characteristics of telescopes which observe electromagnetic radiation other than visible light. Knowing the optimum location of the observatory (ground-based or space-based) and the reasons for the location the observatory are very important.
  - ✓ As a complete understanding of the advantages and limitations of different telescopes, the concepts of aperture, light-gathering power (LGP), angular resolution, focal length, and Dawes Limit should be understood.
  - ✓ With the improved data-handling and analysis capabilities of the late-20<sup>th</sup> and 21<sup>st</sup> centuries, long-baseline radio interferometry and long-baseline optical arrays are becoming more common. The advantages of constructing such arrayed telescopes should be understood.

**DEFINITIONS** You should have a working knowledge of at least these terms and any others used in lecture and lab. Many of these terms will be found in the glossary at the class website.

adaptive optics  
angular resolution  
baseline  
Cassegrain geometry  
Catadioptric design  
charge-coupled device (CCD)

chromatic aberration  
spherical aberration  
coma  
Cassegrain focus  
coudé focus  
Newtonian focus  
prime focus

diffraction  
diffraction grating  
eyepiece  
field  
apparent field  
true field  
focal length

focal plane  
focal point  
focus (as in a lens or  
mirror)  
grating  
imaging  
interferometry  
light-gathering power  
light pollution  
magnification  
magnifying power  
Newtonian reflector  
Cassegrain reflector  
Schmidt-Cassegrain

Maksutov-Cassegrain  
objective  
objective lens  
objective mirror  
primary mirror  
secondary mirror  
optical telescope  
 $\gamma$ -ray telescope  
x-ray telescope  
UV telescope  
IR telescope  
radio telescope  
photometry  
radio window

reflecting telescope  
reflector  
refracting telescope  
refractor  
reflection  
refraction  
seeing  
spectrograph  
spectroscope  
spectroscopy  
very-long-baseline  
interferometry